

# VISUALIZING GEOMETRY WITH THE HELP OF "MATHEMATICA"

VALERY MARENICH AND ANDREY MARENICH

ABSTRACT. Among eight possible geometric structures on three-dimensional manifolds less studied from the differential geometric point of view are those modeled on the Heisenberg group  $Heis^3$  and  $PSL_2(R)$  - the group of all isometries of the Lobachevsky plane, see [Sc]. We consider their left-invariant metrics, define Levi-Civita connections and curvature tensors, define and solve equations of geodesic lines. Using "Mathematica" software package we also present drawings of geodesic lines and metric balls in these spaces.

## 1. Geometry of the Heisenberg group $Heis^3$

**1.1. Left-invariant metric, Levi-Civita connection and curvature tensor of  $Heis^3$ .** We begin with a well-known description of the Heisenberg group of dimension 3. Let  $R^3$  be the euclidean space with coordinates  $(x, y, z)$ . Then the Heisenberg group  $Heis^3$  is this space with the following multiplication rule:

$$(1.1) \quad (x, y, z) \cdot (x', y', z') = (x + x', y + y', z + z' + \langle x, y' \rangle - \langle x', y \rangle),$$

where  $\langle, \rangle$  is a scalar product in  $R^2 = \{(x, y)\}$ . The element zero  $0 = (0, \dots, 0)$  is the unit of this group structure and the vector fields

$$(1.2) \quad X = (1, 0, -y), \quad Y = (0, 1, x), \quad T = (0, 0, 1)$$

are their left-invariant fields. We define the left-invariant metric on  $Heis^3$  by taking  $X, Y, T$  as **the orthonormal frame**.

**Definition 1.** Denote by  $g$  the left-invariant metric on  $Heis^3$  such that vector fields  $X, Y$  and  $T$  are orthonormal ones. The corresponding scalar product we denote as usual by  $(, )$ .

Because due to (1.2) coordinate vectors are

$$\frac{\partial}{\partial x} = (X + yT), \quad \frac{\partial}{\partial y} = (Y - xT), \quad \text{and} \quad \frac{\partial}{\partial z} = T,$$

and by our choice  $\{X, Y, T\}$  is an orthonormal basis we arrive at the following formula for the metric tensor of our left-invariant metric in coordinates  $(x, y, z)$ :

$$(1.3) \quad g = \begin{pmatrix} 1 + y^2 & -xy & y \\ -xy & 1 + x^2 & -x \\ y & -x & 1 \end{pmatrix}.$$

The following was proved in [Mr1].

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1991 *Mathematics Subject Classification.* 53C15, 53C20.

*Key words and phrases.* Left invariant metric, geodesic lines, Heisenberg and  $PSL_2(R)$ -structures.

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**Proposition 1.** *For the covariant derivatives of the Riemannian connection of the left-invariant metric, defined above the following is true:*

$$(1.4) \quad \nabla = \begin{pmatrix} 0 & T & -Y \\ -T & 0 & X \\ -Y & X & 0 \end{pmatrix},$$

where the  $(i, j)$ -element in the table above equals  $\nabla_{E_i} E_j$  for our basis

$$\{E_k, k = 1, 2, 3\} = \{X, Y, T\}.$$

**1.2. Geodesic lines in  $Heis^3$ .** We find equations of geodesics issuing from  $\mathbf{0}=(0,0,0)$ . Let  $c(t)$  be such a geodesics with a natural parameter  $t$ , and its vector of velocity given by

$$(1.5) \quad \dot{c}(t) = \alpha(t)X(t) + \beta(t)Y(t) + \gamma(t)T.$$

Then the equation of a geodesic  $\nabla_{\dot{c}(t)} \dot{c}(t) \equiv 0$  and our table of covariant derivatives (1.3) give:

$$(\alpha'(t) + 2\gamma\beta(t))X(t) + (\beta'(t) - 2\gamma\alpha(t))Y(t) + \gamma'(t)T = 0.$$

Thus we easily obtain the following equations for coordinates of the vector of velocity of the geodesic  $c(t)$  in our left-invariant moving frame:

$$(1.6) \quad \begin{cases} \alpha'(t) + 2\gamma\beta(t) = 0 \\ \beta'(t) - 2\gamma\alpha(t) = 0 \end{cases} \quad \gamma'(t) = 0$$

or

$$(1.7) \quad \begin{cases} (\alpha(t) + \beta(t))' - 2\gamma(\alpha(t) - \beta(t)) = 0 \\ (\alpha(t) - \beta(t))' + 2\gamma(\alpha(t) + \beta(t)) = 0 \end{cases} \quad \gamma'(t) = 0$$

Because the parameter  $t$  is natural we have

$$\alpha^2(t) + \beta^2(t) + \gamma^2 \equiv 1,$$

and we could take  $\gamma(t) \equiv \gamma$  where  $|\gamma| \leq 1$  is the cos of the angle between  $\dot{c}(0)$  and  $T$ -axe. For  $|\gamma| = 1$  we have "vertical" geodesic, coinciding with  $z$ -axe, which is an integral line of the left-invariant vector field  $T$ . For  $\gamma = 0$  our equations are linear. For  $\gamma \neq 0$  after some easy computation one could find that:

$$(1.8) \quad \begin{cases} \alpha(t) = r \cos(2\gamma t + \phi) \\ \beta(t) = r \sin(2\gamma t + \phi) \end{cases}$$

where  $r = \sqrt{\alpha^2 + \beta^2}$ . To find in coordinates  $\{x, y, z\}$  equations for geodesics  $c(t) = (x(t), y(t), z(t))$  issuing from  $\mathbf{0}$  note, that if

$$\dot{c}(t) = \alpha(t)X(t) + \beta(t)Y(t) + \gamma(t)T$$

and our left-invariant vector fields are

$$X = (1, 0, -y), \quad Y = (0, 1, x), \quad T = (0, 0, 1),$$

then

$$\frac{\partial}{\partial x} = X + yT, \quad \text{and} \quad \frac{\partial}{\partial y} = Y - xT,$$

so we easily have:

$$(1.9) \quad \begin{cases} \dot{x}(t) = \alpha(t) \\ \dot{y}(t) = \beta(t) \\ \dot{z}(t) = \gamma - \alpha(t)y(t) + \beta(t)x(t) \end{cases}$$

that after some computations gives the following equations for geodesics issuing from zero:

**Proposition 2.** *Geodesic lines issuing from zero  $\mathbf{0}$  in the Heisenberg group  $Heis^3$  satisfy to the following equations:*

$$(1.10) \quad \begin{cases} x(t) = \frac{r}{2\gamma}(\sin(2\gamma t + \phi) - \sin(\phi)) \\ y(t) = \frac{r}{2\gamma}(\cos(\phi) - \cos(2\gamma t + \phi)) \\ z(t) = \frac{1+\gamma^2}{2\gamma}t - \frac{1-\gamma^2}{4\gamma^2} \sin(2\gamma t) \end{cases}$$

for some numbers  $\phi, r$  which could be defined from the initial condition  $\dot{c}(0) = (r \cos(\phi), r \sin(\phi), \gamma)$ ; or if  $\gamma = 0$ , then they are "horizontal" and satisfy the following equations:

$$(1.11) \quad \begin{cases} x(t) = \alpha(0)t \\ y(t) = \beta(0)t \\ z(t) = 0 \end{cases}$$

To find geodesics issuing from an arbitrary point and their equations it is sufficient to use left translation to this point and apply to the equation above the multiplication rule (1).

### 1.3. Computer generated pictures of geodesic lines and metric balls in $H^3$ .



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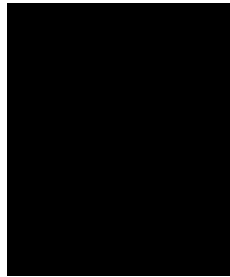
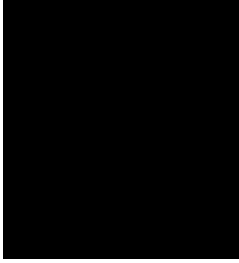
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1. Two geodesic lines issuing from  $\mathbf{0}$  in directions of  $X$  and  $Y$  axes for  $t \in [-10, 10]$ .
2. Exp-image of the  $\{X, T\}$ -coordinate plane.

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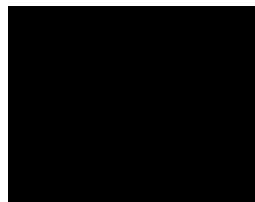
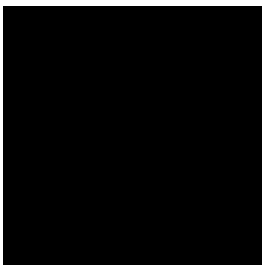


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1. Metric ball with the center at  $\mathbf{0}$  and radius 1.
2. Metric ball with the center at  $\mathbf{0}$  and radius 3.
3. Half of the ball of radius 5.
4. Focal point of the exponential map: singular point of the sphere with radius 5
5. Amplified singular point of the sphere of radius 5.
6. Amplified neighborhood of the singular point of the same sphere.
7. Singular point of the metric sphere of radius  $t = 20$ .

## II. On the geometry of $PSL_2(R)$

### 2.1. Ball model of Lobachevsky plane and global coordinates in $\tilde{P}SL_2(R)$ .

In a standard way we identify the Lobachevsky hyperbolic plane  $H^2$  with the unit disk  $B^2 = \{z \in \mathbb{C} \mid \|z\| < 1\}$  in the complex plane with the metric

$$ds^2 = \frac{dx^2 + dy^2}{(1 - (x^2 + y^2))^2},$$

where  $z = x + iy$ . In this coordinate system isometries can be described as transformations of  $B^2$  of the form

$$z \rightarrow w(z) = -e^{i\phi} \frac{z - \alpha}{\bar{\alpha}z - 1} \quad \text{for } \alpha \in B^2, 0 \leq \phi \leq 2\pi$$

(which is a composition of a parallel translation taking the point 0 to  $\alpha$  and rotation on angle  $\phi$ ). Then  $(\alpha, \phi)$  for  $\alpha \in B^2$  and  $\phi \in S^1 = \{z \in \mathbb{C} \mid \|z\| = 1\}$  are coordinates in the group  $PSL_2(R)$  of all isometries of  $H^2$ . Respectively, in the universal cover  $\tilde{P}SL_2(R)$  of  $PSL_2(R)$  the system of global coordinates is  $(\alpha, \phi)$ , where  $\phi \in \mathbb{R}$ . Below we denote by  $I(\alpha, \phi)$  the isometry with coordinates  $(\alpha, \phi)$  and also  $\alpha = x + iy$ , so that  $(x, y, \phi)$  are our global coordinates in  $\tilde{P}SL_2(R)$ .

### 2.2. Multiplication law in $\tilde{P}SL_2(R)$ .

In the coordinate system which we have introduced above the multiplication law giving coordinates of a composition of two isometries with coordinates  $(\alpha, \phi)$  and  $(\beta, \psi)$  can be defined in the following way. Because

$$I(\alpha, \phi) \circ I(\beta, \psi)(z) = -e^{\phi+\psi} \frac{\gamma}{\bar{\gamma}} \frac{z - \delta}{\bar{\delta}z - 1}$$

for  $\gamma = \gamma(\alpha, \beta) = 1 + \alpha\bar{\beta}e^{-i\psi}$  and  $\delta = \delta(\alpha, \beta) = (\beta + \alpha e^{-i\psi})/\gamma$ , we deduce that

$$I(\alpha, \phi) \circ I(\beta, \psi) = I(\delta, \phi + \psi + 2\text{Arg}(\gamma))$$

which we simply denote as

$$(2.1) \quad (\alpha, \phi) \circ (\beta, \psi) = (\delta(\alpha, \beta), \phi + \psi + 2\text{Arg}(\gamma(\alpha, \beta))).$$

**2.3. Left-invariant vector fields, Lie algebra  $psl_2(R)$  and metric tensor.** Take three coordinate vectors  $X = \partial/\partial x$ ,  $Y = \partial/\partial y$  and  $T = \partial/\partial \phi$  of the coordinate system introduced above at the point  $0 = (0, 0, 0)$ , and applying the differential of the multiplication (2.1) by  $(\alpha, \phi)$  from the left we obtain three left invariant vector fields  $X(\alpha, \phi)$ ,  $Y(\alpha, \phi)$  and  $T(\alpha, \phi)$  correspondingly. As direct calculations show these vector fields in our coordinates  $(x, y, \phi)$  are

$$(2.2) \quad X(x, y, \phi) = (1 - x^2 + y^2, -2xy, 2y), \quad Y(x, y, \phi) = (-2xy, 1 + x^2 - y^2, -2x) \quad \text{and} \quad T(x, y, \phi) = (y, -x, 1).$$

Also direct computations show that

$$[X, Y] = -4T \quad [T, X] = Y \quad [Y, T] = X,$$

or taking from now on  $E_1 = X/2$ ,  $E_2 = Y/2$  and  $E_3 = T$  we have for the left-invariant vector fields  $E_i$ ,  $i = 1, 2, 3$

$$[E_1, E_2] = \lambda_3 E_3 \quad [E_3, E_1] = \lambda_2 E_2 \quad [E_2, E_3] = \lambda_1 E_1$$

where

$$\lambda_1 = \lambda_2 = 1 \quad \text{and} \quad \lambda_3 = -1.$$

**Definition 2.** Denote by  $g$  the left-invariant metric on  $P\tilde{S}L_2(R)$  such that vector fields  $E_i, i = 1, 2, 3$  are orthonormal ones. The corresponding scalar product we denote as usual by  $(, )$ .

Because due to (2.2) coordinate vectors are

$$\frac{\partial}{\partial x} = \frac{1}{1 - (x^2 + y^2)}(2E_1 - 2yE_3), \quad \frac{\partial}{\partial y} = \frac{1}{1 - (x^2 + y^2)}(2E_2 + 2xE_3),$$

and

$$\frac{\partial}{\partial \phi} = E_3 - y\frac{\partial}{\partial x} + x\frac{\partial}{\partial y} = \frac{1}{1 - (x^2 + y^2)}(-2yE_1 + 2xE_2 + (1 + x^2 + y^2)E_3),$$

and by our choice  $\{E_1, E_2, E_3\}$  is an orthonormal basis we arrive at the following formula for the metric tensor of our left-invariant metric in coordinates  $(x, y, \phi)$ :

$$(2.3) \quad g = \frac{1}{(1 - (x^2 + y^2))^2} \begin{pmatrix} 4 + 4y^2 & -4xy & -4y - 2y(1 + x^2 + y^2) \\ -4xy & 4 + 4x^2 & 4x + 2x(1 + x^2 + y^2) \\ -4y - 2y(1 + x^2 + y^2) & 4x + 2x(1 + x^2 + y^2) & (1 + x^2 + y^2)^2 + 4x^2 + 4y^2 \end{pmatrix}.$$

#### 2.4. Curvature tensor of $\tilde{P}SL_2(R)$ .

Applying well known formulas (see [M]) for the curvature tensor of the left-invariant metric  $g$  we obtain the following. Ricci curvature of  $g$ :

$$(2.4) \quad Ric(E_1) = -\frac{3}{2} \quad Ric(E_2) = -\frac{3}{2} \quad Ric(E_3) = \frac{1}{2},$$

scalar curvature

$$(2.5) \quad \rho = -\frac{5}{2},$$

and for sectional curvatures  $K_{ij}$  in two-dimensional directions generated by  $\{E_i, E_j\}$

$$(2.6) \quad K_{12} = -\frac{7}{4} \quad K_{13} = \frac{1}{4} \quad K_{23} = \frac{1}{4}.$$

Note, that there is a natural projection  $\pi : P\tilde{S}L_2(R) \rightarrow B^2$  such that  $\pi(x, y, \phi) = (x, y)$  which is a riemannian submersion (i.e., is an isometry in horizontal directions normal to the fibres of the projection  $\pi$ ). Indeed, by the well-known O'Neill's formula (see [O'N]) we see that the sectional curvature of the base  $H^2$  of this submersion equals

$$(2.7) \quad K(H^2) = K_{12} + \frac{3}{4} \|[E_1, E_2]\|^2 = -\frac{7}{4} + \frac{3}{4} = -1.$$

#### 2.5. Cristoffel symbols of Levi-Civita connection.

**Proposition 3, see [Mr2].** *For the covariant derivatives of the riemannian connection of the left-invariant metric  $g$  defined above the following is true:*

$$(2.8) \quad \nabla = \begin{pmatrix} 0 & -\frac{1}{2}E_3 & 2E_2 \\ \frac{1}{2}E_3 & 0 & 2E_1 \\ 3E_2 & -3E_1 & 0 \end{pmatrix},$$

where the  $(i, j)$ -element in the table above equals  $\nabla_{E_i} E_j$  for our orthonormal basis  $E_i, i = 1, 2, 3$ .

### 2.6. Equation of geodesics.

First we find equations of geodesics issuing from  $\mathbf{0}=(0,0,0)$ . Let  $c(t)$  be such a geodesics with a natural parameter  $t$ , and its vector of velocity is given by

$$(2.9) \quad \dot{c}(t) = u'(t)E_1(c(t)) + v'(t)E_2(c(t)) + w'(t)E_3(c(t)).$$

Then the equation of a geodesic  $\nabla_{\dot{c}(t)}\dot{c}(t) \equiv 0$  and our table of covariant derivatives (2.8) give:

$$(u''(t) - 5v'(t)w'(t))E_1(c(t)) + (v''(t) + 5u'(t)w'(t))E_2(c(t)) + w''(t)E_3(c(t)) = 0,$$

which easily gives

$$(2.10) \quad w'(t) \equiv w'(0)$$

and

$$(2.11) \quad u'(t) = u'(0) \cos(5w'(0)t) + v'(0) \sin(5w'(0)t) \quad \text{and} \quad v'(t) = v'(0) \cos(5w'(0)t) - u'(0) \sin(5w'(0)t).$$

To find equations for the coordinates  $(x(t), y(t), \phi(t))$  of the geodesic line  $c(t)$  in our coordinate system  $(x, y, \phi)$  recall that due to (2.2) we have

$$E_1 = (1 - x^2 + y^2) \frac{\partial}{\partial x} + (-2xy) \frac{\partial}{\partial y} + (2y) \frac{\partial}{\partial \phi}, \quad E_2 = (-2xy) \frac{\partial}{\partial x} + (1 + x^2 - y^2) \frac{\partial}{\partial y} + (-2x) \frac{\partial}{\partial \phi},$$

and

$$(2.12) \quad E_3 = y \frac{\partial}{\partial x} + (-x) \frac{\partial}{\partial y} + \frac{\partial}{\partial \phi}.$$

Therefore, (2.10)-(2.11) substituted in (2.9) gives

$$(2.13) \quad \begin{aligned} \dot{c}(t) = & (u'(t)(1 - x^2 + y^2) + v'(t)(-2xy) + w'(0)y) \frac{\partial}{\partial x} + \\ & (u'(t)(-2xy) + v'(t)(1 + x^2 - y^2) - w'(0)x) \frac{\partial}{\partial y} + (u'(t)(2y) + v'(t)(-2x) + w'(0)) \frac{\partial}{\partial \phi}, \end{aligned}$$

so that from  $\dot{c}(t) = x'(t) \frac{\partial}{\partial x} + y'(t) \frac{\partial}{\partial y} + \phi'(t) \frac{\partial}{\partial \phi}$  and that obviously  $x'(0) = u'(0)$ ,  $y'(0) = v'(0)$  and  $\phi'(0) = w'(0)$  we arrive at the following equations

$$(2.14) \quad \begin{aligned} \dot{x}(t) = & (1 - x^2(t) + y^2(t))(u' \cos(5w't) + v' \sin(5w't)) + (-2x(t)y(t))(v' \cos(5w't) - u' \sin(5w't)) + w'y(t) \\ \dot{y}(t) = & (-2x(t)y(t))(u' \cos(5w't) + v' \sin(5w't)) + (1 + x^2(t) - y^2(t))(v' \cos(5w't) - u' \sin(5w't)) - w'x(t), \\ \dot{\phi}(t) = & (2y(t))(u' \cos(5w't) + v' \sin(5w't)) + (-2x(t))(v' \cos(5w't) - u' \sin(5w't)) + w' \end{aligned}$$

where we drop argument 0 in  $u'(0)$ ,  $v'(0)$  and  $w'(0)$  so that  $(u', v', w')$  equals the unit vector  $\dot{c}(0)$ . First two equations are independent of the third one, hence we first concentrate on the equation for  $\alpha(t) = (x(t) + iy(t))$ . If we denote  $u' \cos(5w't) + v' \sin(5w't) = \sqrt{1 - (w')^2} \cos(\tau)$ , then  $v' \cos(5w't) - u' \sin(5w't) = (\sqrt{1 - (w')^2}) \sin(\tau)$  for  $\tau = 5w't$ , and due to  $x^2 - y^2 = Re(\alpha^2)$  and  $2xy = Im(\alpha^2)$  we find

$$(\dot{x}(t) + i\dot{y}(t)) = \dot{\alpha}(t) =$$

$$\begin{aligned}
& \sqrt{1 - (w')^2}((1 - Re(\alpha^2(t))) \cos(\tau) + (iIm(\alpha^2(t)))(i \sin(\tau) + (-iIm(\alpha^2(t))) \cos(\tau) + i(1 + Re(\alpha^2(t))) \sin(\tau)) \\
& \quad - (iw')(iIm(\alpha(t))) - (iw')Re(\alpha(t)) = \\
& \quad \sqrt{1 - (w')^2}((1 - \alpha^2(t)) \cos(\tau) + (1 + \alpha^2(t))(i \sin(\tau))) - iw'\alpha(t) = \\
(2.15) \quad & \sqrt{1 - (w')^2}(Exp^{i\tau} - \alpha^2(t)Exp^{-i\tau}) - iw'\alpha(t).
\end{aligned}$$

Changing variable  $t$  to  $s = \sqrt{1 - (w')^2}t$  last equality yields for

$$u(s) = \alpha(s/\sqrt{1 - (w')^2})Exp^{i\frac{w's}{\sqrt{1 - (w')^2}}}$$

the following equations

$$\dot{u}(s) = Exp^{isA} - u^2(s)Exp^{-isA}$$

where  $A = 6w'/\sqrt{1 - (w')^2}$ , which in turn gives Ricatti equation

$$\dot{r}(s) = 1 + (iA)r(s) - r^2(s)$$

for  $r(s) = Exp^{iAs}u^{-1}(s)$ . Finally, with the help of software package "Mathematica" we get

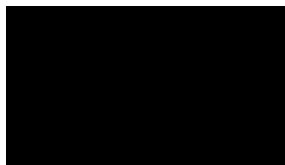
$$(2.16) \quad r(s) = \frac{1}{2}(Ai + (\sqrt{A^2 - 4})tg[-\frac{\sqrt{A^2 - 4}}{2}s + C_1]),$$

where  $C_1 = \pi/2$  due to initial conditions. Finally, we arrive at the following formula for the first two coordinates of the geodesic line:

$$(2.17) \quad \alpha(t) = \frac{Exp^{\frac{5w'}{\sqrt{1 - (w')^2}}s}}{\frac{1}{2}(\frac{6w'i}{\sqrt{1 - (w')^2}} + (\sqrt{\frac{40(w')^2 - 4}{1 - (w')^2}})ctg[(\sqrt{10(w')^2 - 1})t]}.$$

Substituting  $x(t) = Re(\alpha(t))$  and  $y(t) = Im(\alpha(t))$  into the equation (2.14) for the third coordinate, we define  $\phi(t)$ , and with the help of "Mathematica" obtain the following pictures describing geodesic lines and metric balls in  $PSL_2(R)$ .

### 2.7. Computer generated pictures of geodesic lines in $PSL_2(R)$ .





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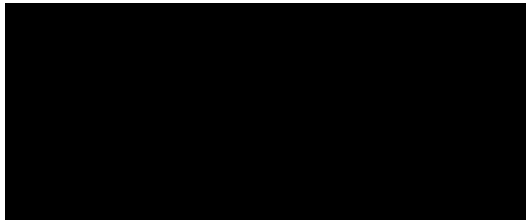
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1. Horizontal projection of the geodesic line issuing from  $\mathbf{0}$  in the direction  $(0.91, 0, 0.3)$ .
2. Horizontal projection of the geodesic line issuing from  $\mathbf{0}$  in the direction  $(0.84, 0, 0.4)$ .
3. Horizontal projection of the geodesic line issuing from  $\mathbf{0}$  in the direction  $(0.19, 0, 0.9)$ .
4. Geodesic line issuing from  $\mathbf{0}$  in the direction  $(0.9975, 0, 0.05)$
5. Geodesic line issuing from  $\mathbf{0}$  in the direction  $(0.99, 0, 0.1)$
6. Geodesic line issuing from  $\mathbf{0}$  in the direction  $(0.91, 0, 0.3)$
7. Geodesic line issuing from  $\mathbf{0}$  in the direction  $(0.64, 0, 0.6)$
8. Metric ball of the radius 0.1.
9. Half of the sphere of the radius 1.
10. Part of the  $\{X, \Phi\}$  coordinate plane in the neighborhood of the focal point.

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